Dark Matter Co-annihilating with a Top/Bottom Partner with effects of bound states

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Background on DM annihilation and decoupling



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Solution by Co-annihilation, [Griest and Seckel].



- If the partner (squark) too heavy, DM relic density overproduced
- ▶ Parameter space: $m_{\chi} = 100 \text{ GeV} \sim 1 \text{ TeV}$, $m_{\tilde{q}} - m_{\chi} < 100 \text{ GeV}$.

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Furthermore, the bound states (onia) formed by co-annihilation partners Y and \overline{Y} is denoted by B.

Sommerfeld effect [Hisano et al., Cirelli et al.] (√s > 2m_Y) in the annihilation of the colored co-annihilating partner pair Y Y
, whose wavefunction at the origin is strongly distorted due to the gluon exchange between the two initial particles. e.g. q
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- $XX \rightarrow SM$ and YY to SM included

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- ▶ Bound state *B* radiative capture and dissociation (labelled by $B \leftrightarrow Y$): $Y\bar{Y} \leftrightarrow Bg$

Physics quantities \mathcal{Y}, z, γ

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 n is the number density of a species and
 s is the entropy density in the universe.
- $z \equiv m_X/T$ with Boltzmann distribution for calculating the thermal number densities. T is the temperature of the universe.

$$\gamma = \int d\Phi_i d\Phi_f(2\pi)^4 \delta^4(\Sigma p) |\mathcal{M}(i \leftrightarrow f)|^2 e^{-(E_i)/T}$$

Coupled Boltzmann equations involving X, Y, B

$$sHz\frac{d\mathcal{Y}_X}{dz} = -\gamma_{XX} \left[\left(\frac{\mathcal{Y}_X}{\mathcal{Y}_X^{eq}} \right)^2 - 1 \right] - 2\gamma_{XY} \left[\frac{\mathcal{Y}_X}{\mathcal{Y}_X^{eq}} \frac{\mathcal{Y}_Y}{\mathcal{Y}_Y^{eq}} - 1 \right] \\ - 2\gamma_{X\leftrightarrow Y} \left[\frac{\mathcal{Y}_X}{\mathcal{Y}_X^{eq}} - \frac{\mathcal{Y}_Y}{\mathcal{Y}_Y^{eq}} \right] - 2\gamma_{B\leftrightarrow X} \left[\left(\frac{\mathcal{Y}_X}{\mathcal{Y}_X^{eq}} \right)^2 - \frac{\mathcal{Y}_B}{\mathcal{Y}_B^{eq}} \right],$$

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$$-\frac{2\gamma_{X\leftrightarrow Y}}{\mathcal{Y}_X^{eq}} \left[\frac{\mathcal{Y}_X}{\mathcal{Y}_X^{eq}} - \frac{\mathcal{Y}_Y}{\mathcal{Y}_Y^{eq}} \right] - \frac{2\gamma_{B\leftrightarrow X}}{\mathcal{Y}_X^{eq}} \left[\left(\frac{\mathcal{Y}_X}{\mathcal{Y}_X^{eq}}\right)^2 - \frac{\mathcal{Y}_B}{\mathcal{Y}_B^{eq}} \right],$$

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$$sHz\frac{d\mathcal{Y}_B}{dz} = -\gamma_{B\leftrightarrow X} \left[\frac{\mathcal{Y}_B}{\mathcal{Y}_B^{eq}} - \left(\frac{\mathcal{Y}_X}{\mathcal{Y}_X^{eq}}\right)^2 \right] - \gamma_{B\leftrightarrow Y} \left[\frac{\mathcal{Y}_B}{\mathcal{Y}_B^{eq}} - \left(\frac{\mathcal{Y}_Y}{\mathcal{Y}_Y^{eq}}\right)^2 \right] - \gamma_{B\leftrightarrow SM} \left[\frac{\mathcal{Y}_B}{\mathcal{Y}_B^{eq}} - 1 \right]$$

•
$$\mathcal{Y}_Y = \mathcal{Y}_{\bar{Y}}$$
, $\gamma_{X\bar{Y}} = \gamma_{XY}$, and $\gamma_{X\leftrightarrow\bar{Y}} = \gamma_{X\leftrightarrow Y}$ by CP.

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$$\therefore \gamma_{B\leftrightarrow X,Y,g} \gg sHz,$$

$$\frac{\mathcal{Y}_B}{\mathcal{Y}_B^{eq}} \simeq \frac{\gamma_{B\leftrightarrow X} \left(\frac{\mathcal{Y}_X}{\mathcal{Y}_X^{eq}}\right)^2 + \gamma_{B\leftrightarrow Y} \left(\frac{\mathcal{Y}_Y}{\mathcal{Y}_Y^{eq}}\right)^2 + \gamma_{B\leftrightarrow SM}}{\gamma_{B\leftrightarrow Y} + \gamma_{B\leftrightarrow X} + \gamma_{B\leftrightarrow SM}}$$

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 $\mathcal{Y}_{\text{dark}} = \mathcal{Y}_X + \mathcal{Y}_Y + \mathcal{Y}_{\bar{Y}}$

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$$\gamma_{\text{bound state}} = \frac{(\gamma_{B\leftrightarrow X} + \gamma_{B\leftrightarrow Y})\gamma_{B\leftrightarrow SM}}{\gamma_{B\leftrightarrow X} + \gamma_{B\leftrightarrow Y} + \gamma_{B\leftrightarrow SM}}$$

Sommerfeld

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$$\sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow hh}v_{rel} \rightarrow S_{1} \times (\sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow hh}v_{rel})|_{v_{rel}\to 0}$$
$$\sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow gh}v_{rel} \rightarrow S_{8} \times (\sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow gh}v_{rel})\Big|_{v_{rel}\to 0}$$

Sommerfeld

$$\begin{split} \mathcal{S}_{1} &= \frac{4\pi\alpha_{3}/(3v)}{1 - \exp[-4\pi\alpha_{3}/(3v)]} > 1 \\ \mathcal{S}_{8} &= \frac{\pi\alpha_{3}/(6v)}{\exp[\pi\alpha_{3}/(6v)] - 1} < 1 \\ \sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow hh}v_{rel} \rightarrow \mathcal{S}_{1} \times (\sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow hh}v_{rel})|_{v_{rel}\to 0} \\ \sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow gh}v_{rel} \rightarrow \mathcal{S}_{8} \times (\sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow gh}v_{rel})\Big|_{v_{rel}\to 0} \\ \hline \sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow gg}v_{rel} \rightarrow \left(\frac{2}{7}\mathcal{S}_{1} + \frac{5}{7}\mathcal{S}_{8}\right) \times (\sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow gg}v_{rel})\Big|_{v_{rel}\to 0} \\ \sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow \bar{q}q}v_{rel} \rightarrow \mathcal{S}_{1} (\sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow (h^{*})\leftrightarrow \bar{q}q}v_{rel})\Big|_{v_{rel}\to 0} + \mathcal{S}_{8} (\sigma_{\tilde{t}\tilde{t}^{*}\leftrightarrow (g^{*})\leftrightarrow \bar{q}q}v_{rel})\Big|_{v_{rel}\to 0} \end{split}$$

QM, Krammer extended



$$(\sigma v_{\rm rel})_{Y\bar{Y}\to Bg}^{n,\ell} = \frac{4\alpha_D}{81\pi}\omega_n \left[\ell \left| \int dr r^3 \left(\omega_n - \Delta V\right) R_{n\ell} R_{k\ell-1} \right|^2 + (\ell+1) \left| \int dr r^3 \left(\omega_n - \Delta V\right) R_{n\ell} R_{k\ell+1} \right|^2 \right]$$

 $\omega_n = E_n + k^2/(2\mu), \ \Delta V(r) = \Delta V_1(r) + \Delta V_2(r),$

$$\Delta V_1(r) = \frac{4\alpha_S^{(f)}}{3r} + \frac{\alpha_S^{(i)}}{6r} , \ \Delta V_2(r) = \frac{3\alpha_S^{(f)}}{2r} .$$



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$$H_{int}^{(1)} = \frac{g_S T_{i'i}^C \delta_{j'j}}{m_Y} \vec{k} \cdot \vec{G}^C(\vec{r}/2) + \frac{g_S \delta_{i'i} T_{j'j}^C}{m_Y} \vec{k} \cdot \vec{G}^C(-\vec{r}/2)$$



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$$\mathcal{V}_{fi}^{(1)} = \frac{\delta^{CD}}{\sqrt{6}} \frac{g_S}{\mu} \langle \Psi_f | \vec{k} | \Psi_i \rangle \cdot \vec{\varepsilon}_{\lambda}^C$$

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$$\vec{k} = -i\mu[\vec{r},\mathsf{KE}] = -i\mu\left(\vec{r}H_8 - H_1\vec{r} - \Delta V_1(r)\vec{r}
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$$\begin{split} \vec{k} &= -i\mu[\vec{r},\mathsf{K}\mathsf{E}] = -i\mu\left(\vec{r}H_8 - H_1\vec{r} - \Delta V_1(r)\vec{r}\right)\\ \hat{H}_1 &= \vec{k}^2/(2\mu) - 4\alpha_s^{(f)}/(3r), \text{ and } \hat{H}_8 = \vec{k}^2/(2\mu) + \alpha_s^{(i)}/(6r). \end{split}$$

$$\mathcal{V}_{fi}^{(1)} = \frac{\delta^{CD}}{\sqrt{6}} (-ig_S) \langle \Psi_f | \left(E_i - E_f - \Delta V_1(r) \right) \vec{r} | \Psi_i \rangle \cdot \vec{\varepsilon}_{\lambda}^C$$



$$\mathcal{L}_{3G} = g_S f^{ABC} (\vec{\nabla} G_0^A) G_0^B \cdot \vec{G}^C$$



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$$\mathcal{M} = 2g_S^3 T_{i'i}^A T_{j'j}^B f^{ABC} \frac{\vec{q} \cdot \vec{\varepsilon}^C}{\vec{q}^4} , \quad H_{int}^{(2)} = -g_S \alpha_S T_{i'i}^A T_{j'j}^B f^{ABC} \vec{r} \cdot \vec{G}^C$$

$$\mathcal{M} = 2(ig_S T^A_{i'i}) \left(\frac{-i}{q^2}\right) (ig_S f^{ABC}) (i\vec{q} \cdot \vec{\varepsilon}^C) \left(\frac{-i}{q^2}\right) (-ig_S T^B_{j'j})$$

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$$\mathcal{V}_{fi}^{(2)} = ig_S \delta^{CD} \sqrt{\frac{1}{6}} \left\langle \Psi_f \left| \underbrace{\frac{3\alpha_S}{2r}}_{\Delta V_2(r)} \vec{r} \right| \Psi_i \right\rangle \cdot \vec{\varepsilon}_{\lambda}^C$$

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Bino-Stop

$\mathcal{L}_{\text{bino-stop}} = \sqrt{2}g_1 \tfrac{2}{3} (\bar{\chi} \tfrac{1}{2} (1+\gamma_5)t) \tilde{t}^* + \text{h.c.}$

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$$\Gamma_{B_{\bar{t}}\to\chi\chi} = \frac{128\pi\alpha_1^2 m_\chi^2 \left(m_{\bar{t}}^2 - m_\chi^2\right)^{3/2}}{27m_{\tilde{t}}^3 \left(m_{\tilde{t}}^2 - m_\chi^2 + m_t^2\right)^2} |\Psi(0)|^2$$
$$\Gamma_{B_{\bar{t}}\to t\bar{t}} = \frac{64\pi\alpha_1^2 m_t^2 \left(m_{\tilde{t}}^2 - m_\chi^2\right)^{3/2}}{81m_{\tilde{t}}^3 \left(m_{\tilde{t}}^2 - m_t^2 + m_\chi^2\right)^2} |\Psi(0)|^2$$

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LHC constraint on the decay $\tilde{t} \to t^* + \chi \to W^* + b + \chi$.

Most complete calculation of DM relic abundance including colored partner states of masses near DM.

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- Attention of the color in the Krammer process for the bound state formation.

Backup material

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