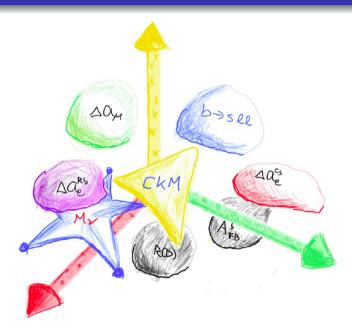
A colorful framework to accommodate the observed anomalies nowadays.



We-Fu Chang
National Tsing Hua University
IOPAS HEP seminar, Jan. 21, 2022

Outlines

- Motivation—Anomalies
- The Model and its solutions
- Constraints and Numerical Study
- Phenomenological consequences
- Origin of the flavor pattern?
- Conclusion

See JHEP 09 (2021) 043 for detailed ref.

Standard Model (SM)

= special relativity +

particle content +

gauge symmetries and breaking

+ QM

QFT

Quarks +Leptons + Higgs

 $SU\left(3\right)_{C}\times SU\left(2\right)_{L}\times U\left(1\right)_{Y}$

gluons W^\pm, Z^0, γ

One of the highest intellectual achievements



Standard Model

One of the highest intellectual achievements

 $\triangle Q_{e}$ electron anomalous magnetic moment

$$\triangle Q_e^{\text{exp}} = 1159652180.73 (28) \times 10^{-12}$$

PRA 83, 052122 (2011)

 $\triangle Q_e^{\text{th}} = 1159652182.032 (720) \times 10^{-12}$

PRD 97, 036001 (2018) $\triangle Q_e^{\text{th}} \ge 0.16$)

g=1 in classical EM g=2 by Dirac Eq. $g=2\left(1+\Delta a\right)$

2

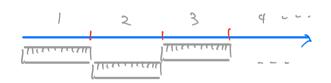
What does 10^{-12} accuracy mean ?





$$1AU = 1.5 \times 10^{11} m$$
$$10^{-12} AU \sim 15 cm$$

Say, each measurement per second



 $1y \sim \pi \times 10^7 \,\mathrm{sec}$

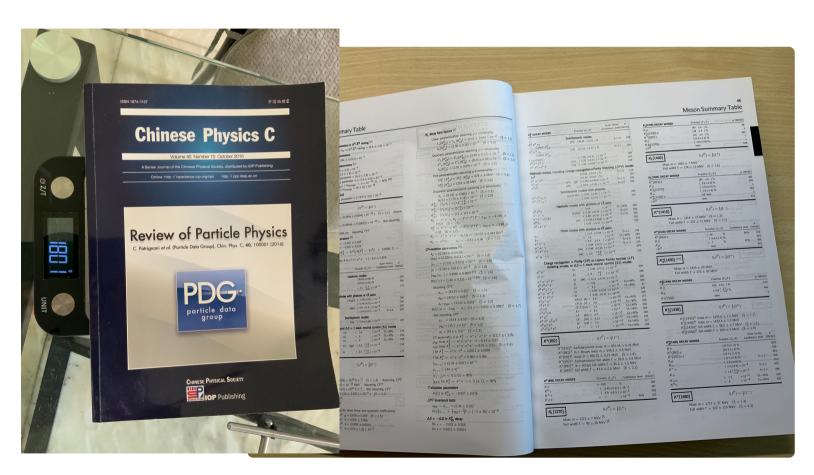
(W/o any rest)

You need to live for $\sim 3 \times 10^4$ years to complete

this task, and ONLY 1 measurement mistake is allowed!!

2,5

Not only that



However, SM is not the end of the story; it is incomplete.

SM=QM+SR QFT

Quantum Gravity? Dark Energy?

+particle content

Quarks +Leptons

+ Higgs

Neutrino mass?

Antimatter? CPV?

Flavor pattern?

+ gauge symmetries and breaking

 $SU(3)_C \times SU(2)_L \times U(1)_Y$

Non-perturbative

 $\begin{cases} \text{confinement} \\ \text{proton mass} \end{cases} \Rightarrow \begin{cases} \text{Models} \\ \text{Lattice} \end{cases}$

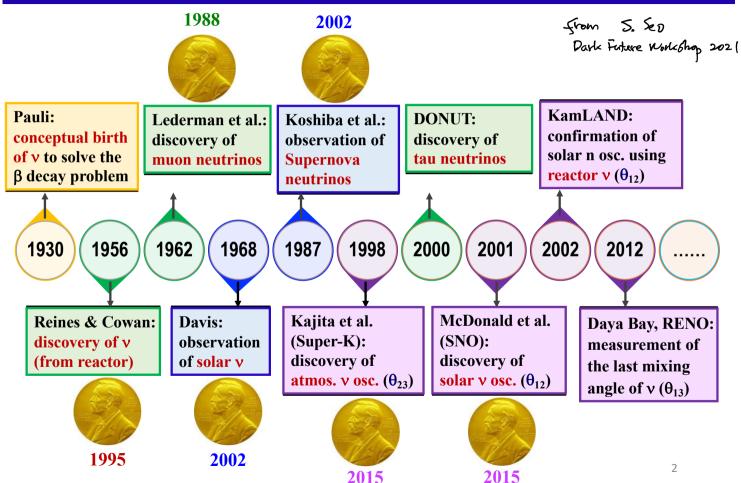
Dark Matter?

Why 3 generations?

We might need some revolutionary ideas

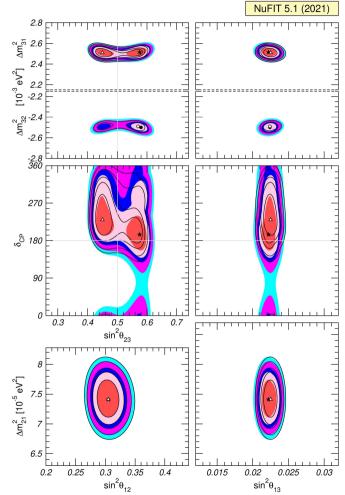
the main playground for the traditional theoretical particle phys.

Milestones of ν History



1/6. Introduction—My Latest nu-data fit (Oct. 2021)

| | | | | | NuFIT 5.1 (2021) | |
|-----------------------------|---------------------------------------------------|---------------------------------|-----------------------------|---------------------------------------------|-----------------------------|--|
| | | Normal Ordering (best fit) | | Inverted Ordering ($\Delta \chi^2 = 2.6$) | | |
| without SK atmospheric data | | bfp $\pm 1\sigma$ | 3σ range | bfp $\pm 1\sigma$ | 3σ range | |
| | $\sin^2 \theta_{12}$ | $0.304^{+0.013}_{-0.012}$ | $0.269 \rightarrow 0.343$ | $0.304^{+0.012}_{-0.012}$ | $0.269 \rightarrow 0.343$ | |
| | $\theta_{12}/^{\circ}$ | $33.44^{+0.77}_{-0.74}$ | $31.27 \rightarrow 35.86$ | $33.45^{+0.77}_{-0.74}$ | $31.27 \rightarrow 35.87$ | |
| | $\sin^2 \theta_{23}$ | $0.573^{+0.018}_{-0.023}$ | $0.405 \rightarrow 0.620$ | $0.578^{+0.017}_{-0.021}$ | $0.410 \rightarrow 0.623$ | |
| | $\theta_{23}/^{\circ}$ | $49.2^{+1.0}_{-1.3}$ | $39.5 \rightarrow 52.0$ | $49.5^{+1.0}_{-1.2}$ | $39.8 \rightarrow 52.1$ | |
| | $\sin^2 \theta_{13}$ | $0.02220^{+0.00068}_{-0.00062}$ | $0.02034 \to 0.02430$ | $0.02238^{+0.00064}_{-0.00062}$ | $0.02053 \to 0.02434$ | |
| | $\theta_{13}/^{\circ}$ | $8.57^{+0.13}_{-0.12}$ | $8.20 \rightarrow 8.97$ | $8.60^{+0.12}_{-0.12}$ | $8.24 \rightarrow 8.98$ | |
| | $\delta_{\mathrm{CP}}/^{\circ}$ | 194^{+52}_{-25} | $105 \rightarrow 405$ | 287^{+27}_{-32} | $192 \rightarrow 361$ | |
| | $\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$ | $7.42^{+0.21}_{-0.20}$ | $6.82 \rightarrow 8.04$ | $7.42^{+0.21}_{-0.20}$ | $6.82 \rightarrow 8.04$ | |
| | $\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$ | $+2.515^{+0.028}_{-0.028}$ | $+2.431 \rightarrow +2.599$ | $-2.498^{+0.028}_{-0.029}$ | $-2.584 \rightarrow -2.413$ | |
| | | Normal Ord | dering (best fit) | Inverted Ordering ($\Delta \chi^2 = 7.0$) | | |
| | | bfp $\pm 1\sigma$ | 3σ range | bfp $\pm 1\sigma$ | 3σ range | |
| | $\sin^2 \theta_{12}$ | $0.304^{+0.012}_{-0.012}$ | $0.269 \rightarrow 0.343$ | $0.304^{+0.013}_{-0.012}$ | $0.269 \rightarrow 0.343$ | |
| lata | $\theta_{12}/^{\circ}$ | $33.45^{+0.77}_{-0.75}$ | $31.27 \rightarrow 35.87$ | $33.45^{+0.78}_{-0.75}$ | $31.27 \rightarrow 35.87$ | |
| ric (| $\sin^2 \theta_{23}$ | $0.450^{+0.019}_{-0.016}$ | $0.408 \rightarrow 0.603$ | $0.570^{+0.016}_{-0.022}$ | $0.410 \rightarrow 0.613$ | |
| with SK atmospheric data | $\theta_{23}/^{\circ}$ | $42.1_{-0.9}^{+1.1}$ | $39.7 \rightarrow 50.9$ | $49.0^{+0.9}_{-1.3}$ | $39.8 \rightarrow 51.6$ | |
| | $\sin^2 \theta_{13}$ | $0.02246^{+0.00062}_{-0.00062}$ | $0.02060 \to 0.02435$ | $0.02241^{+0.00074}_{-0.00062}$ | $0.02055 \to 0.02457$ | |
| | $\theta_{13}/^{\circ}$ | $8.62^{+0.12}_{-0.12}$ | $8.25 \rightarrow 8.98$ | $8.61^{+0.14}_{-0.12}$ | $8.24 \rightarrow 9.02$ | |
| | $\delta_{\mathrm{CP}}/^{\circ}$ | 230^{+36}_{-25} | $144 \rightarrow 350$ | 278^{+22}_{-30} | $194 \rightarrow 345$ | |
| | $\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$ | $7.42^{+0.21}_{-0.20}$ | $6.82 \rightarrow 8.04$ | $7.42^{+0.21}_{-0.20}$ | $6.82 \rightarrow 8.04$ | |
| | $\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$ | $+2.510^{+0.027}_{-0.027}$ | $+2.430 \to +2.593$ | $-2.490^{+0.026}_{-0.028}$ | $-2.574 \rightarrow -2.410$ | |



We have kept seeing many ambulances passing by.





We have kept seeing many ambulances passing by.

Some wrecked · · · ·



We have kept seeing many ambulances passing by.

Some wrecked · · · ·

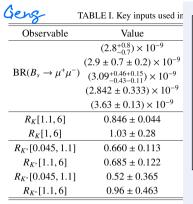
Some keep going . . .

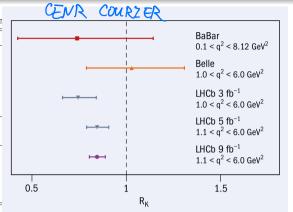


$b \rightarrow sll$

•
$$R_K^{(*)} = \frac{\Gamma\left(B \to K^{(*)}\mu\mu\right)}{\Gamma\left(B \to K^{(*)}ee\right)}$$

$$R_K^{SM} = 1.0004_{-0.0007}^{+0.0008}$$





$$\mathcal{H}_{\text{eff}}^{b \to s\mu\mu} = -\frac{G_F}{\sqrt{2}} \widetilde{V}_{tb} \widetilde{V}_{ts}^* \frac{\alpha}{\pi} \sum_{i} C_i \mathcal{O}_i + H.c. \qquad \mathcal{O}_9 = \left(\bar{s}\gamma^{\alpha} \hat{L}b\right) (\bar{\mu}\gamma_{\alpha}\mu) , \quad \mathcal{O}_{10} = \left(\bar{s}\gamma^{\alpha} \hat{L}b\right) (\bar{\mu}\gamma_{\alpha}\gamma^5\mu) , \\ \mathcal{O}_9' = \left(\bar{s}\gamma^{\alpha} \hat{R}b\right) (\bar{\mu}\gamma_{\alpha}\mu) , \quad \mathcal{O}_{10}' = \left(\bar{s}\gamma^{\alpha} \hat{R}b\right) (\bar{\mu}\gamma_{\alpha}\gamma^5\mu) .$$

Global fit:(after Moriond 2021)

Altmannshofer and Stangl, 2103. 13370

$$C_9 \simeq -0.73$$
 or $C_9 \simeq -C_{10} \simeq -0.39$

(theoretically clean modes only $C_{10} \simeq 0.60\,,~C_9 \simeq -C_{10} \simeq -0.35$)

Geng, Grinstein et al, 2103,12738

$$C_9 \simeq -0.82$$
, $C_9 \simeq -C_{10} \simeq -0.40$
 $C_{10} \simeq 0.65$

$$\Delta a_l$$

$$\overrightarrow{\mu}$$
 (Magnetic moment) $\sim I\overrightarrow{A}$

$$\sim I \overrightarrow{A}$$

by EM

by Dimension-Analysis
$$\left[\overrightarrow{A} \right] = \left(\underbrace{\overrightarrow{QL}}_{T} \right)_{+}$$

and it is a pseudo-vector

 $\sim \text{spin}$

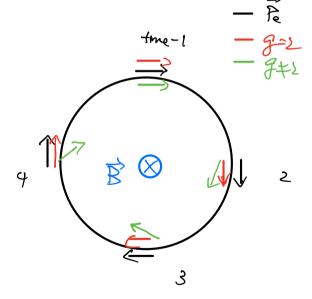
$$\Rightarrow \overrightarrow{\mu} \propto \frac{e}{m} \overrightarrow{S} \qquad \overrightarrow{\mu} = \frac{ge\hbar}{2m} \frac{\overrightarrow{\sigma}}{2}$$

$$\overrightarrow{\mu}=rac{ge\hbar}{2m}rac{\overline{\sigma}}{2}$$

$$g = 1$$
 in classical EM

- Dirac Eq. $\Rightarrow g = 2$ (tree level)
- namely

$$g = 2\left(1 + \Delta a\right)$$



Δa_{l}

$$\Delta a^{(1)} = \frac{\alpha}{2\pi} \sim 1.1617 \times 10^{-3}$$

$$\Delta a^{(2)} \sim O\left(10^{-7}\right)$$

in terms of
$$\mathcal{L} \supset ar{L} \sigma^{\mu \nu} F_{\mu \nu} e_R H$$

$$\Rightarrow \sqrt{m_e} - \zeta$$
 ation $\propto m_e^2$

chirality flip
$$\Rightarrow \propto m_e$$
 another normalization $\propto m_e^2$ $(\text{or } M_F \text{in the loop})$ e^2 $(m_e n_e)$

 $(m_e m_F)$

extremely sensitive to

$$(7 + 3.6) \times 10^{-13}$$

$$\triangle a_e^{Cs} = a_e^{exp} - a_e^{SM} \simeq (-8.7 \pm 3.6) \times 10^{-13}$$

$$\alpha^{-1} = 137.035000046 (27)$$
recall frequency of ¹³

 $\alpha^{-1} = 137.035999046$ (27) , recoil frequency of '33 Cs in matter-wave interferometer 1812.04130

$$\triangle a_e^{Rb} \simeq (+4.8 \pm 3.0) \times 10^{-13}$$
 recoil velocity of 1812 $\alpha^{-1} = 137.035999206 (11)$

Nature 588, 61

Δa_{μ}

$$a_{\mu}^{QED} = 116584718.931 (104) \times 10^{-11}$$

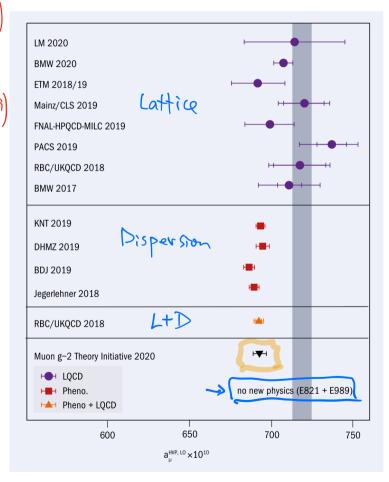
$$a_{\mu}^{EW} = (153.6 \pm 1.0) \times 10^{-11}$$

$$a_{\mu}^{HVP} : \text{ Largest uncertainty} \qquad Q_{\mu}^{EED} \sim O(10^{-7})$$

$$a_{\mu}^{HLBL} \times 10^{11} \sim \begin{cases} 90 \pm 20 & \text{pheno} \\ 80 \pm 35 & \text{cother} \end{cases}$$

$$\Delta a_{\mu} \simeq (25.1 \pm 5.9) \times 10^{-10}$$

CERN COURIER



Cabbibo Angle anomaly

$$V_{CKM} \equiv V_L^u (V_L^d)^\dagger =$$

$$\Rightarrow |V_{nd}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1$$

$$V_{CKM} \equiv V_L^u(V_L^d)^\dagger = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ N_{olg7} & N_{olg7} & N_{olg7} \end{pmatrix} |V_{us}| \sim |V_{cd}| \sim |V_{cd}|$$

$$\Rightarrow \left|V_{nd}\right|^2 + \left|V_{us}\right|^2 + \left|V_{ub}\right|^2 = 1$$
 (Since $\left|V_{ub}\right|^2 \sim 10^{-5}$, $\left|V_{ud}\right|^2 + \left|V_{us}\right|^2 \simeq 1$)

The present PDG value:

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9985 \pm 0.0005$$

From the 2nd neutrino conference(1974), Pennsylvania, (As already pointed out by Y Uchida at CLFV2016.)

• The solar neutrino talk by R.K. Ulrik

AIP Conference Proceedings 22, 259 (1974); doi: 10.1063/1.2947415 The 37 Ar production rates for the standard model and the low Z model are 5.6 \pm 1.8 SNU and 1.4 \pm 0.35 SNU, respectively. Taking Davis's result¹ without run 27 to be 0.2 \pm 0.8 SNU I find that the discrepancy between the experiment and the standard model to be 2.7 σ and while the discrepancy with the low Z model is 1.4 σ .

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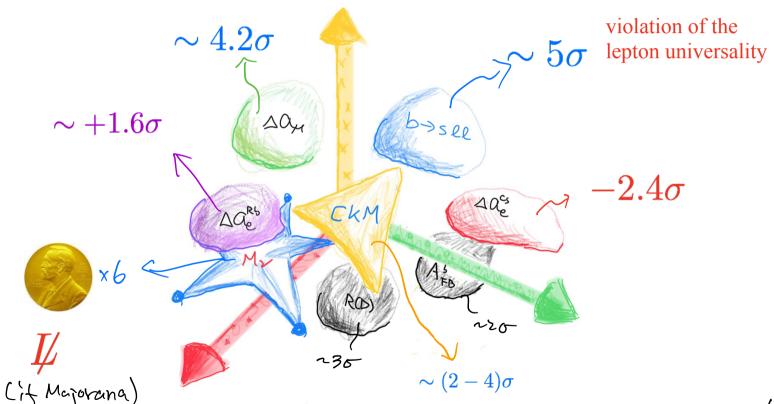
• The Conference Summary by R.P. Feynman

I see no reason to include the solar neutrino problem in here because it is only 1-1/2 standard deviations off of some solar models and the principle is, you don't make a new theory because of 1-1/2 standard deviations.

AIP Conference Proceedings 22, 299 (1974); doi: 10.1063/1.2947418

By statistics, 1 single anomaly might just be the fluctuation.

But, several/many anomalies is more likely the sign of systematics(New phys).



- ① b->s 11: NP connects both lepton/quark sectors
- 2 Neutrino Majorana mass: Lepton# violation
- ③ CKM leakage: mixing between new fermion and the down (or up) sector.
- 4 AMM: implies possible heavy fermion in the loop and the new DOFs must be charged.

What can it be?

A straightforward solution:

- 1 b->s 11: NP connects both lepton/quark sectors
- ② Neutrino Majorana mass: Lepton# violation ⇒ Lo triplet

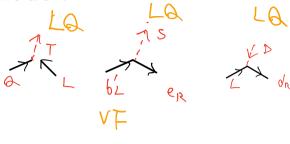
 3 CKM leakage: mixing between new fermion and the
- (3) CKM leakage: mixing between new fermion and the down (or up) sector.
- 4 AMM: implies possible heavy fermion in the loop and the new DOFs must be charged.

La doublet +smglet

-> Leptoguark

• 4 color states are employed in the model:

| | New Fermion | New Scalar | | |
|------------------------|--------------------|---------------------------------------------------------------------------------------------|-------------------------------------------------------------------------|------------------|
| Fields | $b_{L,R}^{\prime}$ | $T = \begin{pmatrix} T^{\frac{2}{3}} \\ T^{-\frac{1}{3}} \\ T^{-\frac{4}{3}} \end{pmatrix}$ | $D = \begin{pmatrix} D^{\frac{2}{3}} \\ D^{-\frac{1}{3}} \end{pmatrix}$ | $S^{rac{2}{3}}$ |
| $SU(3)_c$ | 3 | 3 | 3 | 3 |
| $SU(3)_c$ $SU(2)_L$ | 1 | 3 | 2 | 1 |
| $U(1)_Y$ | $-\frac{1}{3}$ | $-\frac{1}{3}$ | $\frac{1}{6}$ | $\frac{2}{3}$ |
| lepton number | 0 | 1 | -1 | -1 |
| baryon number | $\frac{1}{3}$ | $\frac{1}{3}$ | $\frac{1}{3}$ | $\frac{1}{3}$ |



- Assuming 2-body decays: 6'> Wt, 67, 6H, Mb'> 1,22TeV ATLAS

 CMS

 Mis = 115 TeV

 AS ref. point
- - Two possible scalar cubic interactions:

$$\mathcal{L} \supset \mu_{3} \left\{ H, \tilde{D} \right\} \odot T + \mu_{1} \left[H, D \right] S^{-\frac{2}{3}} + H.c.$$

$$= \mu_{3} \left[H^{+} D^{\frac{1}{3}} T^{-\frac{4}{3}} - \frac{1}{\sqrt{2}} \left(H^{0} D^{\frac{1}{3}} - H^{+} D^{-\frac{2}{3}} \right) T^{-\frac{1}{3}} - H^{0} D^{-\frac{2}{3}} T^{\frac{2}{3}} \right]$$

$$- \mu_{1} \frac{1}{\sqrt{2}} \left(H^{0} D^{\frac{2}{3}} - H^{+} D^{-\frac{1}{3}} \right) S^{-\frac{2}{3}} + H.c.$$



• Assuming U(1)B, the most general Yukawa is:

$$\begin{split} \mathcal{L} \supset -\widetilde{\lambda}_{T} T^{\dagger} \cdot \left\{ \bar{L}^{c}, Q \right\} - \widetilde{\lambda}_{D} \bar{d}_{R} \left[L, D \right] - \widetilde{\lambda}'_{D} \bar{b'}_{R} \left[L, D \right] - \widetilde{\lambda}_{S} \bar{e}_{R} b'_{L} S^{-\frac{2}{3}} - \widetilde{Y}'_{d} \bar{Q} b'_{R} H + H.c. \\ &= -\widetilde{\lambda}_{T} \left[\bar{\nu}^{c} u_{L} T^{-\frac{2}{3}} + (\bar{\nu}^{c} d_{L} + \bar{e}^{c} u_{L}) \frac{T^{\frac{1}{3}}}{\sqrt{2}} + \bar{e}^{c} d_{L} T^{\frac{4}{3}} \right] - \widetilde{Y}'_{d} (\bar{u}_{L} H^{+} + \bar{d}_{L} H^{0}) b'_{R} \\ &- \widetilde{\lambda}_{D} \frac{\bar{d}_{R}}{\sqrt{2}} \left(\nu_{L} D^{-\frac{1}{3}} - e_{L} D^{\frac{2}{3}} \right) - \widetilde{\lambda}'_{D} \frac{\bar{b'}_{R}}{\sqrt{2}} \left(\nu_{L} D^{-\frac{1}{3}} - e_{L} D^{\frac{2}{3}} \right) - \widetilde{\lambda}_{S} \bar{e}_{R} b'_{L} S^{-\frac{2}{3}} + H.c. \end{split}$$

• In addition, new Dirac masses are allowed $\mathcal{L} \supset M_1 \bar{b}_B' b_L' + M_2 \bar{d}_B b_L' + H.c.$

• down-type quark mass matrix becomes 4x4:

$$\mathcal{L} \supset -(\bar{d_R}, \bar{b_R}) \mathcal{M}^d \begin{pmatrix} d_L \\ b_L' \end{pmatrix} + H.c., \ \mathcal{M}^d = \begin{pmatrix} \frac{\widetilde{Y}_d v_0}{\sqrt{2}} & M_2 \\ \frac{\widetilde{Y}_d' v_0}{\sqrt{2}} & M_1 \end{pmatrix}$$

• It can be diagonalized by 4X4 biunitary rotation, and the CC int becomes

$$\mathcal{L} \supset \frac{g_2}{\sqrt{2}} (\bar{u}, \bar{c}, \bar{t}) \gamma^{\alpha} \hat{L} \widetilde{V} \begin{pmatrix} d \\ s \\ b \\ b' \end{pmatrix} W_{\alpha}^{+} + H.c. \qquad \qquad \widetilde{V} = \begin{pmatrix} \widetilde{V}_{ud} \ \widetilde{V}_{us} \ \widetilde{V}_{ub} \ \widetilde{V}_{ub} \ \widetilde{V}_{ub'} \\ \widetilde{V}_{cd} \ \widetilde{V}_{cs} \ \widetilde{V}_{cb} \ \widetilde{V}_{cb'} \\ \widetilde{V}_{td} \ \widetilde{V}_{ts} \ \widetilde{V}_{tb} \ \widetilde{V}_{tb'} \end{pmatrix} \qquad \text{an unpleasant matrix } \mathcal{V} \subset \mathcal{V}_{cd} \ \widetilde{V}_{td} \ \widetilde{V}_{ts} \ \widetilde{V}_{tb} \ \widetilde{V}_{tb'} \end{pmatrix}$$



• Mathematical trick: consider an auxiliary 4x4 unitary matrix

$$V_4$$

$$\widetilde{V}_{4} \equiv \begin{pmatrix} U_{L}^{u} & 0 \\ 0 & 1 \end{pmatrix} \cdot (U_{L}^{d})^{\dagger} = \begin{pmatrix} \widetilde{V}_{ud} & \widetilde{V}_{us} & \widetilde{V}_{ub} & \widetilde{V}_{ub'} \\ \widetilde{V}_{cd} & \widetilde{V}_{cs} & \widetilde{V}_{cb} & \widetilde{V}_{cb'} \\ \widetilde{V}_{td} & \widetilde{V}_{ts} & \widetilde{V}_{tb} & \widetilde{V}_{tb'} \\ (U_{L}^{d})_{d4}^{*} & (U_{L}^{d})_{s4}^{*} & (U_{L}^{d})_{b4}^{*} & (U_{L}^{d})_{b'4}^{*} \end{pmatrix}$$

For simplicity, assume NO extra CPV

=> 3 extra mixings between [d]-b'

$$(U_L^d)^{\dagger} = \begin{pmatrix} (U_{L3}^d)^{\dagger} & 0 \\ 0 & 1 \end{pmatrix} \cdot R_4, \text{ where } R_4 = \begin{pmatrix} c_1 & 0 & 0 & s_1 \\ -s_1 s_2 & c_2 & 0 & c_1 s_2 \\ -s_1 c_2 s_3 & -s_2 s_3 & c_3 & c_1 c_2 s_3 \\ -s_1 c_2 c_3 & -s_2 c_3 & -s_3 & c_1 c_2 c_3 \end{pmatrix}$$

Finally, a delightful expression for \mathcal{F}_{α}

$$\widetilde{V}_4 = \begin{pmatrix} V_{CKM} & 0 \\ 0 & 1 \end{pmatrix} \cdot R_4$$

$$|\widetilde{V}_{ud}|^2 + |\widetilde{V}_{us}|^2 + |\widetilde{V}_{ub}|^2 = 1 - |\widetilde{V}_{ub'}|^2 \le 1$$



$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9985 \pm 0.0005$$

and

$$\widetilde{V}_{ub'} \simeq 0.9740s_1 + 0.2265c_1s_2 + 0.0036c_1c_2s_3e^{1.196i}$$
.

d-6, s-6'

we have

$$|s_1 + 0.233s_2| \simeq 0.039(7)$$

(note: only 01, 02)

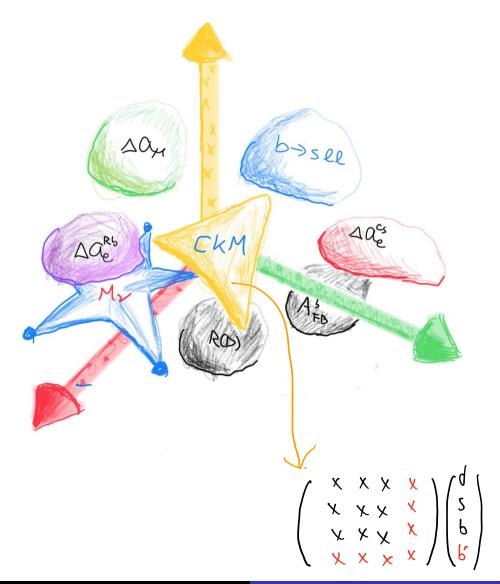
• Other predictions of CKM triangles :

$$\tilde{V}_{ud}\tilde{V}_{ub}^* + \tilde{V}_{cd}\tilde{V}_{cb}^* + \tilde{V}_{td}\tilde{V}_{tb}^* = -(U_L^d)_{d4}^*(U_L^d)_{b4}$$

$$\begin{split} |\widetilde{V}_{cd}|^2 + |\widetilde{V}_{cs}|^2 + |\widetilde{V}_{cb}|^2 &= 1 - |\widetilde{V}_{cb'}|^2, \\ |\widetilde{V}_{td}|^2 + |\widetilde{V}_{ts}|^2 + |\widetilde{V}_{tb}|^2 &= 1 - |\widetilde{V}_{tb'}|^2, \\ |\widetilde{V}_{ud}|^2 + |\widetilde{V}_{cd}|^2 + |\widetilde{V}_{td}|^2 &= 1 - |(U_L^d)_{d4}|^2, \\ |\widetilde{V}_{us}|^2 + |\widetilde{V}_{cs}|^2 + |\widetilde{V}_{ts}|^2 &= 1 - |(U_L^d)_{s4}|^2, \\ |\widetilde{V}_{ub}|^2 + |\widetilde{V}_{cb}|^2 + |\widetilde{V}_{tb}|^2 &= 1 - |(U_L^d)_{b4}|^2, \end{split}$$

• In the mass basis, JL and UL have diff LH rotation.

using charged lepton/down quark as indices



Neutrino Mass

• Consider a general coupling:

$$\mathcal{L} \supset \lambda_{ij} \bar{F}_j \nu_{Li} \phi_1 + \kappa_{ij} \bar{F}_j \nu_i^c \phi_2 + H.c.$$

$$\downarrow \qquad \qquad \downarrow \qquad \downarrow \qquad \qquad \qquad \qquad \downarrow \qquad \qquad \qquad \downarrow \qquad \qquad \downarrow \qquad \qquad \downarrow \qquad \qquad \downarrow \qquad \qquad$$

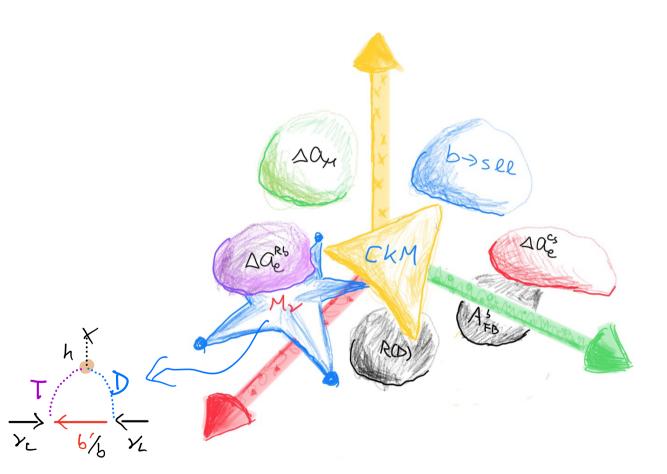
• The 1-loop result is finite and exact:

$$M_{ij}^{\nu} = \sum_{k} \frac{N_c^F m_k}{16\pi^2} s_{\alpha} c_{\alpha} (\kappa_{ik} \lambda_{jk} + \kappa_{jk} \lambda_{ik}) \left[\frac{m_h^2}{m_h^2 - m_k^2} \ln \frac{m_h^2}{m_k^2} - \frac{m_l^2}{m_l^2 - m_k^2} \ln \frac{m_l^2}{m_k^2} \right]$$

∠: the mixing angle between ♦, ≥, ♦

In this model

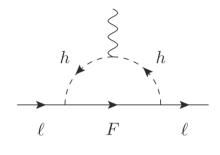
$$M_{ij}^{
u} \simeq \sum_{k=d,s,b,b'} \frac{3m_k}{32\pi^2} (\lambda_{ik}^T \lambda_{jk}^D + \lambda_{jk}^T \lambda_{ik}^D) \frac{\mu_3 v_0}{M_D^2 - M_T^2} \ln \frac{M_T^2}{M_D^2}$$

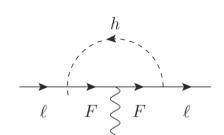




A general coupling

$$\mathcal{L} \supset \bar{F}(y_R^l \hat{R} + y_L^l \hat{L}) \ell h + H.c.$$





 $\left(\Delta a_e^{cs} \Delta a_{\mu} < o\right)$

$$-N_c^F(1+Q_F)m_l^2 f^1$$

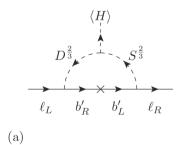
$$\Delta a_l^F = \frac{-N_c^F Q_F m_l^2}{8\pi^2} \int_0^1 dx \, x^2 \frac{(1-x)\frac{|y_L^l|^2 + |y_R^l|^2}{2} + \frac{m_F}{m_l} \Re[(y_R^l)^* y_L^l]}{x^2 m_l^2 + x(m_F^2 - m_l^2) + (1-x)m_h^2} \,,$$

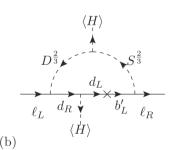
If mock MF,

$$\triangle a_l = \triangle a_l^F + \triangle a_l^h \simeq -\frac{N_c^F \Re[(y_R^l)^* y_L]}{8\pi^2} \left(\frac{m_l}{m_F}\right) \mathcal{J}_{Q_F} \left(\frac{m_h^2}{m_F^2}\right)$$

an d

$$\mathcal{J}_Q(\alpha) = \int_0^1 dx \, \frac{x(1-x) + x \, Q}{x + (1-x)\alpha}$$
$$= \frac{2Q(1-\alpha)(1-\alpha + \alpha \ln \alpha) + (1-\alpha^2 + 2\alpha \ln \alpha)}{2(1-\alpha)^3}$$



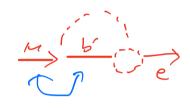




transitione

NOTE: The diagrams also induce the FCNC

A simple arrangement to avoid that,



$$\Rightarrow$$

以 Plays many roles in the model.

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$$Sol-1:$$

$$\triangle a_e \simeq 2.28 \times 10^{-5} \times \left[\lambda_{eb}^D \lambda_{eb}^S\right] \times \left(\frac{\mu_1}{\text{GeV}}\right) \times \mathcal{K}(b_D, b_S),$$

$$\triangle a_{\mu} \simeq 1.03 \times 10^{-10} \times$$

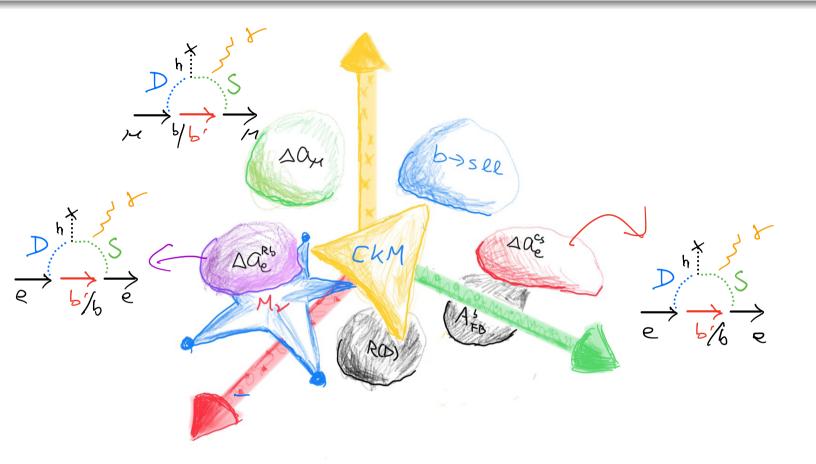
$$\triangle a_e \cong 2.28 \times 10^{-10} \times [\lambda_{eb}^D \lambda_{eb}] \times \left(\overline{\text{GeV}} \right) \times \mathcal{K}(b_D, b_S),$$

$$\triangle a_\mu \simeq 1.03 \times 10^{-10} \times [\lambda_{\mu b'}^D \lambda_{\mu b'}^S] \times \left(\frac{\mu_1}{\text{GeV}} \right) \left(\frac{1.5 \text{TeV}}{M_{b'}} \right)^3 \times \mathcal{K}(\beta_D, \beta_S).$$

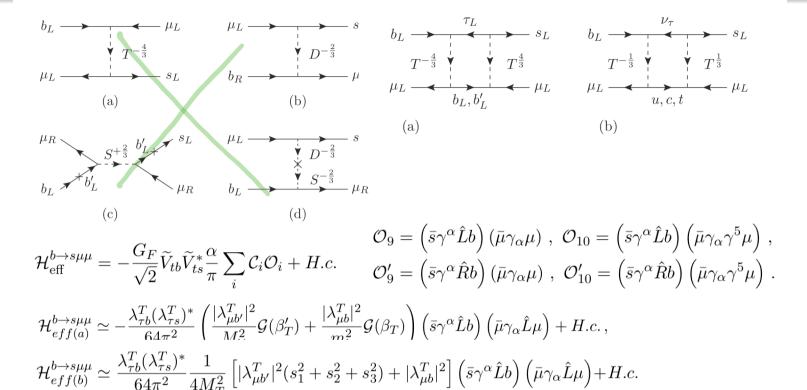
$$Sol-2$$
:

 $\triangle a_e \simeq 5.00 \times 10^{-13} \times [\lambda_{eb'}^D \lambda_{eb'}^S] \times \left(\frac{\mu_1}{\text{GeV}}\right) \left(\frac{1.5 \text{TeV}}{M_V}\right)^3 \times \mathcal{K}(\beta_D, \beta_S),$ $\triangle a_{\mu} \simeq 4.71 \times 10^{-3} \times [\lambda_{\mu b}^{D} \lambda_{\mu b}^{S}] \times \left(\frac{\mu_{1}}{\text{GeV}}\right) \times \mathcal{K}(b_{D}, b_{S}).$

$$\mathcal{K}(a,b) \equiv \frac{\mathcal{J}_{-\frac{1}{3}}(a) - \mathcal{J}_{-\frac{1}{3}}(b)}{b-a}.$$







The box diagram function:
$$G(x) = \left[\frac{1}{1-x} + \frac{\ln x}{(1-x)^2}\right]$$

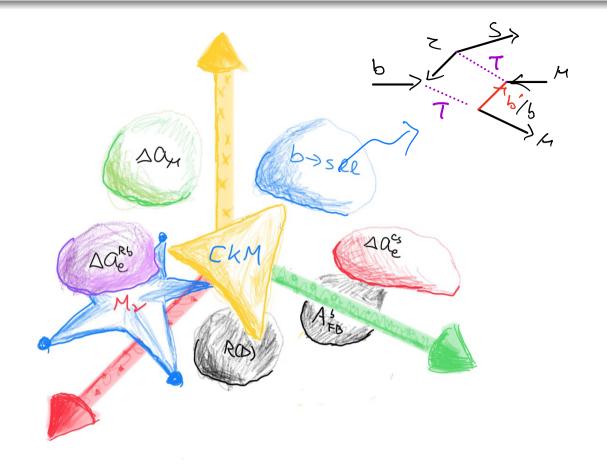
$$C_{9} = -C_{10} \simeq \frac{\sqrt{2}}{128\pi\alpha} \frac{\lambda_{\tau b}^{T} (\lambda_{\tau s}^{T})^{*}}{V_{tb} V_{ts}^{*} G_{F} M_{T}^{2}} \left[|\lambda_{\mu b'}^{T}|^{2} \beta_{T}' \mathcal{G}(\beta_{T}') - \frac{5}{4} |\lambda_{\mu b}^{T}|^{2} \right]$$

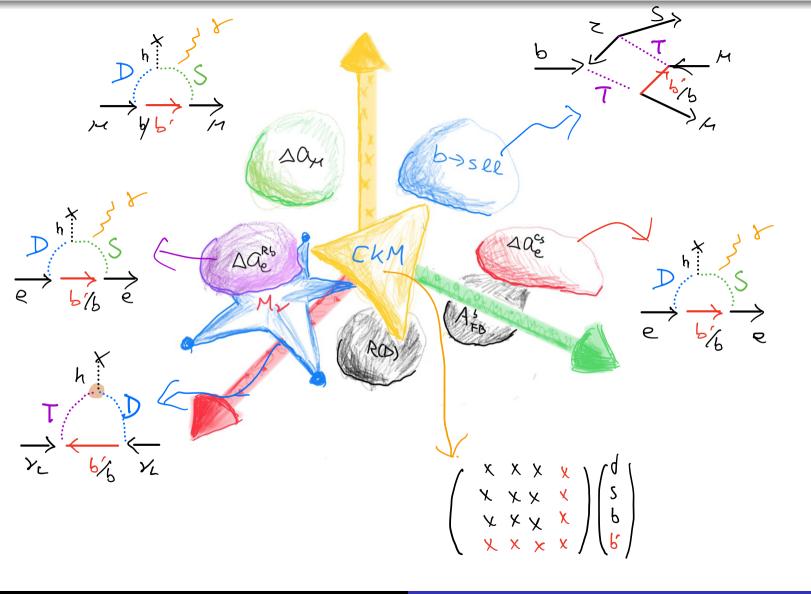
$$C_{9}' = -C_{10}' = 0.$$

And we adopt the best fit values from Altmannshofer& Stangl 2103.13370

$$(\mathcal{C}_9)^{\mu} = -(\mathcal{C}_{10})^{\mu} \simeq -0.41 \pm 0.07,$$

 $(\mathcal{C}_9)^e \simeq (\mathcal{C}_{10})^e \simeq (\mathcal{C}_9')^e \simeq (\mathcal{C}_{10}')^e \simeq 0$





2/6. The Model and its solutions

A Recap

To accommodate the anomalies,

| Anomaly | Requirement |
|-----------------------------------------------------------|------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|
| $m_ u$ | $\mu_3 m_{b^{(\prime)}} \lambda^D \lambda^T \simeq \mathcal{O}(10^{-5}) (\text{GeV})^2$ |
| $\triangle a_e^{Cs[Rb]}, \triangle a_\mu \text{ (Sol-1)}$ | $\left\{ \mu_1 \lambda_{eb}^D \lambda_{eb}^S, \ \mu_1 \lambda_{\mu b'}^D \lambda_{\mu b'}^S \right\} \simeq \{ (49 \pm 20)[-27 \pm 17], (280 \pm 66) \} \text{GeV}$ |
| $\triangle a_e^{Cs[Rb]}, \triangle a_\mu \text{ (Sol-2)}$ | $\left\{ \mu_1 \lambda_{eb'}^D \lambda_{eb'}^S, \ \mu_1 \lambda_{\mu b}^D \lambda_{\mu b}^S \right\} \simeq \{ -(20.1 \pm 8.3)[+11.1 \pm 6.9], -(689 \pm 162) \} \text{GeV}$ |
| $b \to s l^+ l^-$ | $\lambda_{\tau b}^{T}(\lambda_{\tau s}^{T})^{*}(\lambda_{\mu b'}^{T} ^{2} + 3.39 \lambda_{\mu b}^{T} ^{2}) \simeq -(1.07 \pm 0.18)$ |
| Cabibbo angle anomaly | $ s_1 + 0.233s_2 \simeq 0.039(7)$ |

The minimal set of

$$\operatorname{MinS}_{T} = \{ \lambda_{\tau b}^{T}, \ \lambda_{\tau s}^{T}, \ \lambda_{\tau b'}^{T}, \ \lambda_{\mu b'}^{T}, \ \lambda_{\mu b}^{T} \}$$

Need to check the exp. constraints

- * 37 ll operators
- . SM Z° couplings
- · Bs-B, mixing
- $B \rightarrow k^{(k)} \nu \bar{\nu}$
- · 2 > 48, e8
- · 6 >> 5}
- · Neutrino data
- · Orpp decay

- . SM Z° couplings
- · Bs-B, mixing
- $B \rightarrow k^{(*)} V \bar{V}$
- · 2 > 48, e8
- 6->58
- · Neutrino data
- · Oxpp decay

$$\operatorname{MinS}_{T} = \{ \lambda_{\tau b}^{T}, \ \lambda_{\tau s}^{T}, \ \lambda_{\tau b'}^{T}, \ \lambda_{\mu b'}^{T}, \ \lambda_{\mu b}^{T} \}$$

Low energy [qqll] effective operators

| $\frac{(\bar{q}_k \gamma^{\mu} \hat{L} q_l)(\bar{e}_i \gamma_{\mu} \hat{L} e_j)}{4M_T^2}$ | Wilson Coef. | Constraint | Model |
|-------------------------------------------------------------------------------------------|---------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|-------------------------|--------|
| $bb\mu\mu$ | $2 \lambda_{ub}^T ^2$ | 211.1 [116] | 1.06 |
| $sb\tau\tau$ | $2\lambda_{	au b}^{T}(\lambda_{	au s}^{T})^{*}$ | - | -0.14 |
| $sb\mu\mu$ | 0 ^a | - | 0 |
| $sb\mu\tau$ | 0 | - | 0 |
| $sb\tau\mu$ | $2\lambda_{\mu b}^T (\lambda_{	au s}^T)^*$ | 0.199^{b} [116] | 0.11 |
| $uu\tau\mu$ | $\left(\widetilde{V}_{ub}(\lambda_{\tau b}^T)^* + \widetilde{V}_{ub'}(\lambda_{\tau b'}^T)^* + \widetilde{V}_{us}(\lambda_{\tau s}^T)^*\right) \times \left(\widetilde{V}_{ub}^* \lambda_{\mu b}^T + \widetilde{V}_{ub'}^* \lambda_{\mu b'}^T\right]$ | 0.13 [116] | 0.0043 |
| $uu\mu\mu$ | $\left \widetilde{V}_{ub}(\lambda_{\mu b}^T)^* + \widetilde{V}_{ub'}(\lambda_{\mu b'}^T)^*\right ^2$ | 1.03 [116] | 0.017 |
| исµµ | $\left(\widetilde{V}_{ub}(\lambda_{\mu b}^T)^* + \widetilde{V}_{ub'}(\lambda_{\mu b'}^T)^*\right) \times \left(\widetilde{V}_{cb}^*\lambda_{\mu b}^T + \widetilde{V}_{cb'}^*\lambda_{\mu b'}^T\right)$ | 0.11 ^c [116] | 0* |
| сеµµ | $\frac{\left(\widetilde{V}_{ub}(\lambda_{\mu b}^{T})^{*} + \widetilde{V}_{ub'}(\lambda_{\mu b'}^{T})^{*}\right) \times \left(\widetilde{V}_{cb}^{*}\lambda_{\mu b}^{T} + \widetilde{V}_{cb'}^{*}\lambda_{\mu b'}^{T}\right)}{\left \widetilde{V}_{cb}(\lambda_{\mu b}^{T})^{*} + \widetilde{V}_{cb'}(\lambda_{\mu b'}^{T})^{*}\right ^{2}}$ | 52.8 [116] | 0* |
| $cc\tau\mu$ | $\left(\widetilde{V}_{cb}(\lambda_{\tau b}^T)^* + \widetilde{V}_{cb'}(\lambda_{\tau b'}^T)^* + \widetilde{V}_{cs}(\lambda_{\tau s}^T)^*\right) \times \left(\widetilde{V}_{cb}^*\lambda_{\mu b}^T + \widetilde{V}_{cb'}^*\lambda_{\mu b'}^T\right)$ | 211.1 [116] | 0* |
| $tc\mu\mu$ | $\left(\widetilde{V}_{tb}(\lambda_{\mu b}^T)^* + \widetilde{V}_{tb'}(\lambda_{\mu b'}^T)^*\right) \times \left(\widetilde{V}_{cb}^*\lambda_{\mu b}^T + \widetilde{V}_{cb'}^*\lambda_{\mu b'}^T\right)$ | - | 0* |
| $tc\tau\tau$ | $\left(\widetilde{V}_{tb}(\lambda_{\tau b}^T)^* + \widetilde{V}_{tb'}(\lambda_{\tau b'}^T)^* + \widetilde{V}_{ts}(\lambda_{\tau s}^T)^*\right) \times \left(\widetilde{V}_{cb}^* \lambda_{\tau b}^T + \widetilde{V}_{cb'}^* \lambda_{\tau b'}^T + \widetilde{V}_{cs}^* \lambda_{\tau s}^T\right)$ | - | -0.030 |
| $tc\tau\mu$ | $\left(\widetilde{V}_{tb}(\lambda_{\tau b}^T)^* + \widetilde{V}_{tb'}(\lambda_{\tau b'}^T)^* + \widetilde{V}_{ts}(\lambda_{\tau s}^T)^*\right) \times \left(\widetilde{V}_{cb}^*\lambda_{\mu b}^T + \widetilde{V}_{cb'}^*\lambda_{\mu b'}^T\right)$ | 11.35^{d} | 0* |
| $tc\mu\tau$ | $\left(\widetilde{V}_{tb}(\lambda_{\mu b}^T)^* + \widetilde{V}_{tb'}(\lambda_{\mu b'}^T)^*\right) \times \left(\widetilde{V}_{cb}^*\lambda_{\tau b}^T + \widetilde{V}_{cb'}^*\lambda_{\tau b'}^T + \widetilde{V}_{cs}^*\lambda_{\tau s}^T\right)$ | 11.35 | 0.02 |
| $tu\mu\mu$ | $\left(\widetilde{V}_{tb}(\lambda_{\mu b}^T)^* + \widetilde{V}_{tb'}(\lambda_{\mu b'}^T)^*\right) \times \left(\widetilde{V}_{ub}^*\lambda_{\mu b}^T + \widetilde{V}_{ub'}^*\lambda_{\mu b'}^T\right)$ | - | 0.09 |
| $tu\tau\tau$ | $\left(\widetilde{V}_{tb}(\lambda_{\tau b}^T)^* + \widetilde{V}_{tb'}(\lambda_{\tau b'}^T)^* + \widetilde{V}_{ts}(\lambda_{\tau s}^T)^*\right) \times \left(\widetilde{V}_{ub}^* \lambda_{\tau b}^T + \widetilde{V}_{ub'}^* \lambda_{\tau b'}^T + \widetilde{V}_{us}^* \lambda_{\tau s}^T\right)$ | - | -0.03 |
| $tu\tau\mu$ | $\left(\widetilde{V}_{tb}(\lambda_{\tau b}^{T})^{*} + \widetilde{V}_{tb'}(\lambda_{\tau b'}^{T})^{*} + \widetilde{V}_{ts}(\lambda_{\tau s}^{T})^{*}\right) \times \left(\widetilde{V}_{ub}^{*}\lambda_{\mu b}^{T} + \widetilde{V}_{ub'}^{*}\lambda_{\mu b'}^{T}\right)$ | 11.35 | -0.12 |
| $tu\mu\tau$ | $\left(\widetilde{V}_{tb}(\lambda_{\mu b}^T)^* + \widetilde{V}_{tb'}(\lambda_{\mu b'}^T)^*\right) \times \left(\widetilde{V}_{ub}^*\lambda_{\tau b}^T + \widetilde{V}_{ub'}^*\lambda_{\tau b'}^T + \widetilde{V}_{us}^*\lambda_{\tau s}^T\right)$ | 11.35 | 0.02 |

^aThere is no such effective operator at tree-level.

| $\frac{(\bar{d}_k \gamma^{\mu} \hat{L} u_l)(\bar{\nu}_i \gamma_{\mu} \hat{L} e_j)}{4M_T^2}$ | Wilson Coef. | Constraint | Model |
|---------------------------------------------------------------------------------------------|-----------------------------------------------------------------------------------------------------------------------------------------------------------------------|------------|-------------|
| $su\nu_{\mu}\mu$ | 0 | | 0 |
| $su\nu_{\tau}\mu$ | $(\lambda_{\tau s}^T)^* \left(\widetilde{V}_{ub}^* \lambda_{\mu b}^T + \widetilde{V}_{ub'}^* \lambda_{\mu b'}^T \right)$ | 3.96 | 0.010 |
| $su\nu_{\tau}\tau$ | $(\lambda_{\tau s}^T)^* \left(\widetilde{V}_{ub}^* \lambda_{\tau b}^T + \widetilde{V}_{ub'}^* \lambda_{\tau b'}^T + \widetilde{V}_{us}^* \lambda_{\tau s}^T \right)$ | 0.79 | 0.003 |
| $su\nu_{\mu}\tau$ | 0 | | 0 |
| $sc\nu_{\mu}\mu$ | 0 | | 0 |
| $sc\nu_{\tau}\mu$ | $(\lambda_{	au s}^T)^* \left(\widetilde{V}_{cb}^* \lambda_{\mu b}^T + \widetilde{V}_{cb'}^* \lambda_{\mu b'}^T \right)$ | 31.7 | 0* |
| $sc\nu_{\tau}\tau$ | $(\lambda_{\tau s}^T)^* \left(\widetilde{V}_{cb}^* \lambda_{\tau b}^T + \widetilde{V}_{cb'}^* \lambda_{\tau b'}^T + \widetilde{V}_{cs}^* \lambda_{\tau s}^T \right)$ | 15.8 | 0.002 |
| $sc\nu_{\mu}\tau$ | 0 | | 0 |
| $bu\nu_{\mu}\mu$ | $(\lambda_{\mu b}^T)^* \left(\widetilde{V}_{ub}^* \lambda_{\mu b}^T + \widetilde{V}_{ub'}^* \lambda_{\mu b'}^T \right)$ | 0.51 | 0.09 |
| $bu\nu_{\tau}\mu$ | $(\lambda_{\tau b}^T)^* \left(\widetilde{V}_{ub}^* \lambda_{\mu b}^T + \widetilde{V}_{ub'}^* \lambda_{\mu b'}^T \right)$ | 0.51 | -0.12 |
| $bu\nu_{\tau}\tau$ | $(\lambda_{\tau b}^T)^* \left(\widetilde{V}_{ub}^* \lambda_{\tau b}^T + \widetilde{V}_{ub'}^* \lambda_{\tau b'}^T + \widetilde{V}_{us}^* \lambda_{\tau s}^T \right)$ | 0.51 | -0.03 |
| $bu\nu_{\mu}\tau$ | $(\lambda_{\mu b}^T)^* \left(\widetilde{V}_{ub}^* \lambda_{\tau b}^T + \widetilde{V}_{ub'}^* \lambda_{\tau b'}^T + \widetilde{V}_{us}^* \lambda_{\tau s}^T \right)$ | 0.51 | 0.02 |
| $bc\nu_{\mu}\mu$ | $(\lambda_{\mu b}^T)^* \left(\widetilde{V}_{cb}^* \lambda_{\mu b}^T + \widetilde{V}_{cb'}^* \lambda_{\mu b'}^T \right)$ | 5.41 | 0* |
| $bc\nu_{\tau}\mu$ | $(\lambda_{\tau b}^T)^* \left(\widetilde{V}_{cb}^* \lambda_{\mu b}^T + \widetilde{V}_{cb'}^* \lambda_{\mu b'}^T \right)$ | 5.41 | 0* |
| $bc\nu_{	au}	au$ | $(\lambda_{\tau b}^T)^* \left(\widetilde{V}_{cb}^* \lambda_{\tau b}^T + \widetilde{V}_{cb'}^* \lambda_{\tau b'}^T + \widetilde{V}_{cs}^* \lambda_{\tau s}^T \right)$ | 5.41 | -0.03^{a} |
| $bc\nu_{\mu}\tau$ | $(\lambda_{ub}^T)^* \left(\widetilde{V}_{cb}^* \lambda_{\tau b}^T + \widetilde{V}_{cb'}^* \lambda_{\tau b'}^T + \widetilde{V}_{cs}^* \lambda_{\tau s}^T \right)$ | 5.41 | 0.02 |

"This is the effective operator to address the $R(D^{(*)})$ anomaly.

Carpentier Davidson, EPJC70,1071



 $b \rightarrow c \ell_{\lambda} V$ ($\lambda = e.A$) universality better than $\leq 1\%$ (1801.01112)



NO electron counterpart!!

^bWe update this value by using the new data $\mathcal{B}(B^+ \to K^+ \mu^+ \tau^-) < 4.5 \times 10^{-5}$ [1].

^cWe update this value by using the new data $\mathcal{B}(D^+ \to \pi^+ \mu^+ \mu^-) < 7.3 \times 10^{-8}$ [1].

^dWe obtain the limit by using the top quark decay width, $\Gamma_t = 1.42 \text{GeV}$, and $\mathcal{B}(t \to q l l') < 1.86 \times 1.86 \times 1.00 \times 1.$ 10^{-5} [117].

• In the interaction basis

$$\left(g^{sm} - T_3 - \Theta \sin^2\theta w\right),$$

$$\mathcal{L} \supset \frac{g_2}{c_W} \left[g_L^{SM} \sum_{i=1}^3 \bar{d}_{Li} \gamma^\alpha d_{Li} + g_R^{SM} \sum_{i=1}^3 \bar{d}_{Ri} \gamma^\alpha d_{Ri} + g_R^{SM} (\bar{b}_L' \gamma^\alpha b_L' + \bar{b}_R' \gamma^\alpha b_R') \right] Z_\alpha$$

In SM
$$\frac{2b}{R} = \frac{2b}{L} + \frac{1}{2}$$

$$\mathcal{L} \supset \frac{g_2}{c_W} \left[g_L^{SM} \sum_{i=1}^4 \bar{d}_{Li} \gamma^\alpha d_{Li} + g_R^{SM} \sum_{i=1}^4 \bar{d}_{Ri} \gamma^\alpha d_{Ri} + \frac{1}{2} (\bar{b}_L' \gamma^\alpha b_L') \right] Z_\alpha$$

Then, in the mass basis,

$$\int \frac{g_2}{c_W} \left[\sum_{\alpha=s,d,b,b'} \bar{d}_{\alpha} \gamma^{\alpha} (g_L^{SM} \hat{L} + g_R^{SM} \hat{R}) d_{\alpha} \right] Z_{\mu} + \frac{g_2}{2c_W} \sum_{\alpha,\beta=s,d,b,b'} \kappa_{\alpha\beta} \left[(\bar{d}_{\alpha} \gamma^{\alpha} \hat{L} d_{\beta}) \right] Z_{\alpha}$$

$$\kappa_{\alpha\beta} \equiv (U_L^d)_{\alpha4} [(U_L^d)_{\beta4}]^*$$

In our parameterization, the FCNC couplings

$$\kappa_{sd} = \kappa_{ds} = s_1 s_2 c_2 c_3^2$$
, $\kappa_{sb} = \kappa_{bs} = s_2 s_3 c_3$, $\kappa_{bd} = \kappa_{db} = s_1 s_3 c_2 c_3$

- To avoid Tree-level FCNC mediated by ≥°, [ONLY 1 0 ±0]
- For that particular $\partial_{\mathcal{F}}$

$$g_{d_F,R}^{SM} \Rightarrow g_{d_i,R}^{SM}, \ g_{d_F,L}^{SM} \Rightarrow g_{d_iL}^{SM} + \frac{1}{2} \left| (U_L^d)_{4d_F} \right|^2$$

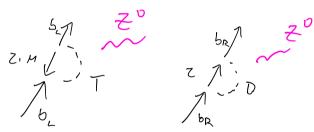
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$$A_F,A_{FB} \propto rac{g_L^2-g_R^2}{g_L^2+g_R^2}\,,\; \Gamma \propto g_L^2+g_R^2\,.$$

 $\overline{g_L^p, A_{FB}} \sim \overline{g_L^2 + g_R^2}, \quad \text{if } S_R^p \sim 0.077$ $\bullet \text{ Long standing LEP puzzle: } A_{FB}^b \sim -2.3\delta \qquad \text{, requires } \theta_3 \neq 0, \Leftrightarrow \text{ckm}$

Only $\theta_1 \neq 0$ is viable to satisfy all Exp. limits

$$|s_1| \simeq 0.039(7)$$
 hegligible δA_{FB}^d



915M~-01423

* How about loop?

In MS and on-shell renormalization

$$\begin{split} \delta g_L^b &\simeq \frac{|\lambda_{\tau b}^T|^2 + |\lambda_{\mu b}^T|^2}{64\pi^2} \left[\left(-1 + \frac{5}{3} s_W^2 \right) \frac{1}{9\beta_Z} - s_W^2 \frac{2 \ln \beta_Z + \frac{1}{3} + i\pi/2}{3\beta_Z} \right] \,, \\ \delta g_R^b &\simeq \frac{|\lambda_{\tau b}^D|^2}{64\pi^2} \left[\left(-\frac{1}{3} s_W^2 \right) \frac{1}{9\beta_Z} + s_W^2 \frac{2 \ln \beta_Z + \frac{1}{3} + i\pi/2}{3\beta_Z} \right] \,, \end{split}$$

Wait for future Z-pole exp. (FCC-ee, ILC, CEPC....)

Neutrino data

For
$$\mathcal{M}_{\gamma}$$
, $\mathcal{M}_{\epsilon b'}$ more parameters are needed
$$\left(\begin{array}{c} \operatorname{MinS}_{T} = \{\lambda_{\tau b}^{T}, \lambda_{\tau b'}^{T}, \lambda_{\mu b'}^{T}, \lambda_{\mu b}^{T}, \lambda_{\mu b}^{T}, \lambda_{\mu b}^{T}, \lambda_{\mu b'}^{T}, \lambda_{\mu b}^{T}, \lambda_{\mu b}^{T},$$

• $M_{\nu}^{ee} = 0$, only NH is allowed

$$\rho_b = m_b/M_{b'} \text{ and } N^{\nu} = \frac{3\mu_3 v_0 M_{b'}}{32\pi^2 M_{LQ}^2}$$

$$\mathcal{M}^{\nu} \simeq \begin{pmatrix} 0 & 0.90792 & 0.13812 \\ 0.90792 & -2.4923 & -2.7643 \\ 0.13812 & -2.7643 & -1.9353 \end{pmatrix} \times 10^{-2} \text{eV} \qquad \begin{pmatrix} \theta_{12} \simeq 33.0^{\circ} \,, & \theta_{23} \simeq 48.7^{\circ} \,, & \theta_{13} \simeq 8.6^{\circ} \,, & \delta_{CP} = 0^{\circ} \,, \\ \triangle m_{21}^{2} \sim 7.47 \times 10^{-5} \text{eV}^{2} \,, & \triangle m_{31}^{2} \sim 2.53 \times 10^{-3} \text{eV}^{2} \,. \end{pmatrix}$$

v.s. 10 best fit (w. SK atm) NuFit

$$\theta_{12} \in (32.7 - 34.21)^{\circ}, \quad \theta_{23} \in (48.0 - 50.1)^{\circ}, \quad \theta_{13} \in (8.45 - 8.69)^{\circ}, \quad \delta_{CP} \in (173 - 224)^{\circ},$$

 $\triangle m_{21}^2 \in (7.22 - 7.63) \times 10^{-5} \text{eV}^2, \qquad \triangle m_{31}^2 \in (2.489 - 2.543) \times 10^{-3} \text{eV}^2.$



The neutrino mass pattern

$$\Delta a_e^{cs} \left[\Delta a_e^{Rb} \right], \Delta a_{\mu},$$
 the Cabibbo angle, b \rightarrow sll anomalies

can be accommodated by the following viable model parameters:

$$\begin{split} M_{LQ} &\simeq 1.0 \, \text{TeV} \,, & M_{b'} = 1.5 \, \text{TeV} \,, & \mu_1 = 2.3 [-0.82] \, \text{TeV} \,, & \mu_3 \simeq 0.5 \, \text{keV} \,, \\ \mu_1 \lambda_{eb'}^S \lambda_{eb'}^D &= -12 [4] \, \text{GeV} \,, & \mu_1 \lambda_{\mu b}^S \lambda_{\mu b}^D &= -690 \, \text{GeV} \,, \\ \theta_2 &= \theta_3 = 0 \,, & \sin \theta_1 = 0.039 \,, \\ \lambda_{\mu b'}^T &\simeq -3.3 \,, & \lambda_{\tau b'}^T \simeq -0.51 \,, & \lambda_{\mu b}^T \simeq -0.72 \,, & \lambda_{\tau b}^T \simeq 0.93 \,, & \lambda_{\tau s}^T \simeq -0.08 \,, \\ \lambda_{eb'}^D &\simeq -0.002 \,, & \lambda_{\mu b}^D \simeq 3.4 \,, & \lambda_{\tau b'}^D \simeq 0.008 \,, & \lambda_{\tau b}^D \simeq -0.58 \,, \end{split}$$



Note that
$$\mathcal{O}\left(\frac{\mu_3}{M_{IO}}\right) \sim 10^{-9}$$
 global L#.

4/6. Phenomenological consequences

stringently set the parameter space boundaries.

b うらて at tree-level

$$\mathcal{H}_{\text{eff}}^{b \to s \tau \tau} \simeq -\frac{G_F}{\sqrt{2}} \tilde{V}_{tb} \tilde{V}_{ts}^* \frac{\alpha}{\pi} \mathcal{C}^{b s \tau \tau} \left[\bar{s} \gamma^{\alpha} \hat{L} b \right] \left[\bar{\tau} \gamma_{\alpha} (1 - \gamma^5) \tau \right] + H.c. ,$$

$$\mathcal{C}^{b s \tau \tau} \simeq \frac{\sqrt{2} \pi}{4 \alpha} \frac{\lambda_{\tau b}^T (\lambda_{\tau s}^T)^*}{\tilde{V}_{tb}^* \tilde{V}_{ts}^* G_F M_T^2} = 23.2 \times \left(\frac{\text{TeV}}{M_T} \right)^2 , \quad \text{C.s.} \quad \text{C$$

$$\Rightarrow B_{\Upsilon}(B_{5} \rightarrow z^{+}z^{-}) \simeq 10^{-5} \qquad \text{SHII} < C \sim O(10^{-3}) \qquad B_{5} \rightarrow z^{+}z^{-} \quad \text{LHCb} \left(1705,02508\right)$$

$$\mathcal{H}_{\text{eff}}^{CC} \supset -\left[\frac{\lambda_{\tau s}^{T}(\lambda_{\tau b}^{T})^{*}\tilde{V}_{cs} + \lambda_{\tau b}^{T}(\lambda_{\tau b'}^{T})^{*}\tilde{V}_{cb'} + |\lambda_{\tau b}^{T}|^{2}\tilde{V}_{cb}}{4M_{T}^{2}}\right] \left(\bar{c}\gamma^{\alpha}\hat{L}b\right) \left(\bar{\tau}\gamma_{\alpha}\hat{L}\nu_{\tau}\right) + H.c.$$



For the benchmark point,

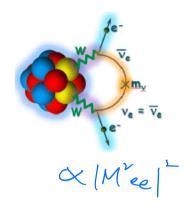
$$\begin{split} \mathcal{B}(T^{-\frac{1}{3}} \to b\nu_{\mu}) &\simeq 18.9\% \,, \quad \mathcal{B}(T^{-\frac{1}{3}} \to b\nu_{\tau}) \simeq 30.9\% \,, \quad \mathcal{B}(T^{-\frac{1}{3}} \to s\nu_{\tau}) \simeq 2.1 \times 10^{-3} \,, \\ \mathcal{B}(T^{-\frac{1}{3}} \to \tau t) &\simeq 30.8\% \,, \quad \mathcal{B}(T^{-\frac{1}{3}} \to \mu t) \simeq 18.9\% \,, \quad \mathcal{B}(T^{-\frac{1}{3}} \to \tau c) \simeq 2.0 \times 10^{-3} \,. \\ \mathcal{B}(T^{\frac{2}{3}} \to t\nu_{\mu}) &\simeq 37.9\% \,, \qquad \mathcal{B}(T^{\frac{2}{3}} \to t\nu_{\tau}) \simeq 61.7\% \,, \qquad \mathcal{B}(T^{\frac{2}{3}} \to c\nu_{\tau}) \simeq 0.4\% \,, \\ \mathcal{B}(T^{-\frac{4}{3}} \to b\mu^{-}) &\simeq 37.9\% \,, \qquad \mathcal{B}(T^{-\frac{4}{3}} \to b\tau^{-}) \simeq 61.7\% \,, \qquad \mathcal{B}(T^{-\frac{4}{3}} \to s\tau^{-}) \simeq 0.4\% \,. \\ \mathcal{B}(D^{-\frac{1}{3}} \to b\bar{\nu}_{\mu}) &\simeq 97.1\% \,, \qquad \mathcal{B}(D^{-\frac{1}{3}} \to b\bar{\nu}_{\tau}) \simeq 2.9\% \,, \\ \mathcal{B}(D^{\frac{2}{3}} \to b\mu^{+}) &\simeq 97.1\% \,, \qquad \mathcal{B}(D^{\frac{2}{3}} \to b\tau^{+}) \simeq 2.9\% \,, \\ \mathcal{B}(S^{\frac{2}{3}} \to b\mu^{+}) &\simeq 100\% \,, \qquad \mathcal{B}(b' \to cW^{-}) \simeq 6 \times 10^{-4} \,, \quad \mathcal{B}(b' \to tW^{-}) \simeq 9 \times 10^{-7} \,, \\ \mathcal{B}(b' \to \bar{\nu}T^{-\frac{1}{3}}) &\simeq 19.7\% \,, \qquad \mathcal{B}(b' \to eS^{\frac{2}{3}}) \simeq 38.5\% \,, \\ \mathcal{B}(b' \to \tau^{+}T^{-\frac{4}{3}}) &\simeq 0.9\% \,, \qquad \mathcal{B}(b' \to eS^{\frac{2}{3}}) \simeq 39.7\% \,, \end{split}$$

Very diff. from the simple assumptions implemented in searching for LQ, b'. (cf. 100% $6 \times 100\%$

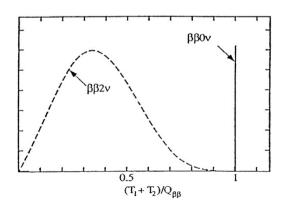
4/6. Phenomenological consequences

Table 2 $T_{1/2}^{0\nu}$ and $\langle m_{\beta\beta} \rangle$ limits (90% CL) from the most recent measurements, sorted by mass number

| Isotope | $T_{1/2}^{0\nu} (\times 10^{25} \text{ years})$ | $\langle m_{\beta\beta} \rangle \text{ (eV)}$ | Experiment | Reference |
|-------------------|-------------------------------------------------|-----------------------------------------------|---------------------|-----------|
| ⁴⁸ Ca | $> 5.8 \times 10^{-3}$ | <3.5-22 | ELEGANT-IV | 159 |
| ⁷⁶ Ge | >8.0 | <0.12-0.26 | GERDA | 160 |
| | >1.9 | <0.24-0.52 | Majorana | 161 |
| | | | Demonstrator | |
| ⁸² Se | $>3.6 \times 10^{-2}$ | <0.89-2.43 | NEMO-3 | 162 |
| ⁹⁶ Zr | $>9.2 \times 10^{-4}$ | <7.2-19.5 | NEMO-3 | 163 |
| $^{100}{ m Mo}$ | $>1.1 \times 10^{-1}$ | < 0.33 - 0.62 | NEMO-3 | 164 |
| ¹¹⁶ Cd | $>2.2 \times 10^{-2}$ | <1.0-1.7 | Aurora | 165 |
| ¹²⁸ Te | $>1.1 \times 10^{-2}$ | NE | C. Arnaboldi et al. | 166 |
| ¹³⁰ Te | >1.5 | < 0.11-0.52 | CUORE | 126 |
| ¹³⁶ Xe | >10.7 | < 0.061 – 0.165 | KamLAND-Zen | 167 |
| | >1.8 | <0.15-0.40 | EXO-200 | 168 |
| ¹⁵⁰ Nd | $>2.0 \times 10^{-3}$ | <1.6-5.3 | NEMO-3 | 169 |



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$$Q_{pp} = M(A,Z) - M(A,Z+2)$$
Due to M-e conv. even with λT_{es} , $\neq 0$

$$|M_{ee}| \leq 3 \times 10^{-4} \text{ eV}$$

- · We need a very special flavor pattern (usually call for some northiveal flavor symmetry)
- In 4D QFT, the Higgs Yukawa after SSB

In 5D, with extra-dimension
 ₹

$$Y_{1L}^{5D} = Y_{1L}^{4D}(X^{h}) + (z) + Y_{2R}^{5D} = Y_{2R}^{4D}(X^{c}) + Q(z)$$

$$H^{5D} = H^{4D}(X^{h}) + (z) + Z_{5D} - Y_{5D} + Y_{5D}^{5D} + Y_{5D}^{$$

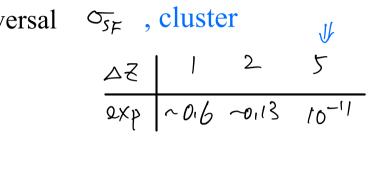
• Integrating out the 5-th dim.

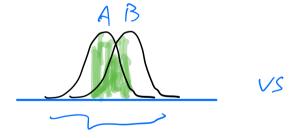
33

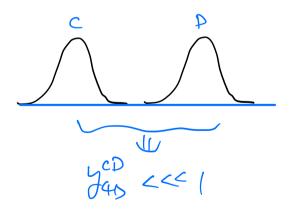
One toy model: split fermion (Arkani-Hamed, 9903417)

All \$\(\phi(\fi) \square \text{Gaussian with universal}\)

$$y_{4D} \sim \exp\left[-\frac{\Delta z^2}{2\sigma_{SF}^2}\right],$$







One example SF configuration

```
\{N_T, N_D, N_S\} = \{4.5, 6.2, 5.5\}.
\{z_{\tau_L}, z_{\mu_L}, z_{e_L}, z_{\mu_R}, z_{e_R}\} = \{0, -0.49, 7.75[7.80], -4.99[-4.72], -0.89[-0.97]\},
\{z_{d_L}, z_{s_L}, z_{b_L}, z_{b'_L}, z_{b_R}, z_{b'_R}\} = \{-5.27[-5.53], 2.86[2.90], -2.14[-2.13],
-1.90[-1.95], -2.06[-2.09], 3.66[3.74]\},
```

which is able to reproduce every parameter within $\sim 50\%$

$$\begin{split} |\lambda_{\mu b'}^T| &= 1.66[1.56]\,, \qquad |\lambda_{\tau b'}^T| = 0.74[0.68]\,, \qquad |\lambda_{\mu b}^T| = 1.15[1.17]\,, \\ |\lambda_{\tau b}^T| &= 0.46[0.46]\,, \qquad |\lambda_{\tau s}^T| = 7.5[6.8] \times 10^{-2}\,, \\ |\lambda_{eb'}^D| &= 1.5[1.6] \times 10^{-3}\,, \quad |\lambda_{\mu b}^D| = 1.78[1.71]\,, \qquad |\lambda_{\tau b'}^D| = 7.5[5.8] \times 10^{-3}\,, \quad |\lambda_{\tau b}^D| = 0.74[0.69]\,, \\ |\lambda_{eb'}^S| &= 3.31[3.42]\,, \qquad |\lambda_{\mu b}^S| = 9.4[19.2] \times 10^{-2}\,. \end{split}$$

and

$$\begin{split} |\lambda_{eb}^{D}| &= 7.5[3.6] \times 10^{-21} \,, \qquad |\lambda_{\mu b'}^{D}| = 1.1[0.8] \times 10^{-3} \,, \\ |\lambda_{ed}^{T}| &= 7.0[0.13] \times 10^{-37} \,, \qquad |\lambda_{es}^{T}| = 2.9[2.8] \times 10^{-5} \,, \qquad |\lambda_{eb}^{T}| = 2.6[1.8] \times 10^{-21} \,, \\ |\lambda_{eb'}^{T}| &= 2.7[1.1] \times 10^{-20} \,, \qquad |\lambda_{\mu d}^{T}| = 4.8[1.4] \times 10^{-5} \,, \qquad |\lambda_{\tau d}^{T}| = 4.2[1.0] \times 10^{-6} \,, \end{split}$$

In addition to the hierarchical Yukawa, the nearly conserved L#, or

$$\mathcal{O}\left(rac{\mu_3}{M_{LQ}}
ight) \sim 10^{-9}$$

can be easily arranged in this 5D model.

• Recall that,
$$\mu_3 \sim 0.5 \text{KeV}$$
, M_{LA} , $M_{b'}$, $\mu_1 \sim O(\text{TeV})$

$$\mathcal{L} > \mu_3 \{H, \tilde{D}\} \circ T + \mu_1 [H, D] S^{-\frac{1}{3}}$$

have the same ... D&S

Overlapping
$$\sim \mathcal{O}(1)$$

e.g. T:
$$(+,-)$$
, D, S: $(-,+)$ on the orbifold

6/6. Conclusion

- Neutrino mass, $b \to sll$, $\triangle a_{\mu}$, $\triangle a_{e}^{Cs[Rb]}$, and CKM anomalies can be accommodated in SM+ 3 scalar leptoquarks (3,3,-1/3),(3,2,1/6),(3,1,2/3) +vector b' with $U(1)_{B}$.
- Viable benchmark point with minimal set of real parameters.
- Solid prediction: normal hierarchy with $\mathcal{M}^{\nu}_{ee} \lesssim 3 \times 10^{-4}$ eV.
- Nontrivial leptoquark b' decay branching ratios.
- Nearly conserved global Lepton number
- Split fermion is one possible origin of the flavor pattern

