### Reactor neutrino spectrum

#### H. B. Li (李浩斌)

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# **Motivation**

- Accurate reactor  $\gamma$  oscillations experiments.
- Better sesitivities on v magnetic search.
- · Reactor monitoring.



SONGS: replacement of Pu was seen. Most of the material based on arXiv:1101.2663v2 [hep-ex]. 
$$\begin{split} & \text{Spectrum}(E) = S_{\text{fission}} + S_{n \text{ capture on }^{238}\text{U}} + S_{n \text{ capture on fission product}} \\ & \text{S}_{\text{fission}}\text{: fission of }^{235}\text{U}, \\ & \text{S}_{n \text{ capture on }}^{238}\text{U}\text{: }^{238}\text{U} + n \rightarrow \\ & \overset{239}{}\text{U} \xrightarrow{} \\ & \overset{239}{}\text{Np} \xrightarrow{} \\ & \overset{239}{}\text{Np} \xrightarrow{} \\ & \overset{239}{}\text{Pu} \\ & \text{S}_{n \text{ capture on fission product}\text{: e. g. }} \\ \end{split}$$

$$S_{\text{fission}}(t,E) = \sum_{k=^{235}\text{U},\;^{238}\text{U},\;^{239}\text{Pu},\;^{241}\text{Pu}} \alpha_k(t) S_k(t,E)$$

 $\alpha_k$ : fission rate, depend on abundance of isotope and neutrons. S<sub>k</sub>:  $\bar{\nu_e}$  or  $e^-$  spectrum per fission.

$$S_k(t, E) = \sum_{fp=\text{fission product}} A_{fp}(t) S_{fp}(E)$$

 $A_{fp}$ : activity of  $fp^{th}$  fission product, depend on neutrons.  $S_{fp}$ :  $\bar{\nu_e}$  or  $e^-$  spectrum of  $fp^{th}$  fission product.

$$S_{fp} = \sum_{b=decay branch} BR_b S_b$$

BR<sub>b</sub>: branching ratio of each decay branch. S<sub>b</sub>:  $\bar{v_e}$  or  $e^-$  spectrum of each  $\beta$ -decay Number of nuclei involve: 845 + unknown. Number of decay branch: >10000 + unknown. Example: A fission of  $^{235}\text{U}$  into  $^{92}\text{Kr}$  and  $^{141}\text{Ba},$  and  $\beta\text{-decay branchs of }^{92}\text{Kr}$ 





 $e^-$  spectrum of <sup>235</sup>U, <sup>239</sup>Pu, <sup>241</sup>Pu are measured at ILL(Institut Laue-Langevin) High-Flux reactor by neutron bombardment on <sup>235</sup>U, <sup>239</sup>Pu, <sup>241</sup>Pu thin foil.





### Convert $e^-$ spectrum to $\bar{\nu_e}$ spectrum: old way



Cut  $e^-$  spectrum into n-bin, the highest E bin must come from largest branch(one branch), assume constant Z

- $\rightarrow$  fit the last bin with SINGLE branch  $e^-$  spectrum
- ightarrow subtract that spectrum from measured  $e^-$  spectrum
- $\rightarrow$  fit the last bin after subtraction.

Z dependent of Fermi function affect jigsaw structure of low Energy  $\bar{v_e}$  spectrum.

# Add all the fission product and $\beta$ -branch

A near complete data available at ENSDF(Evaluated Nuclear Structure Data File).

However ...



 $E_0$  and branching ratio was measured by  $\gamma$  spectrum, and  $\gamma$  could "lost" in measurement  $\rightarrow$  assign larger  $E_0$  $\rightarrow$  Pandemonium Effect(Hardy, 1977) Using Total Absorption Gamma Spectrometer(TAGS).

## Compare with Measured $e^-$ spectrum



ENSDF only, replace some with Pandemonium-corrected data, add in JENDL(Japanese Evaluated Nuclear Data Library) and model.

 $\pm 10\%$ 

### Another way

Add everythings in ENSDF and Pandemonium-corrected data, and fit remaining as "old way".



The remaning are fitted with 5 virtual branches with Z=46.

 $e^-$  spectrum  $\pm 1\%$ 

 $\bar{\nu_e}$  spectrum shift +3%



Use ENSDF only to generated  $\bar{\nu_e}$  and  $e^-$ , then convert generated- $e^-$  spectrum to  $\bar{\nu_e}$  with old way.



Switch on-off various effects  $\rightarrow$  +3% below 4 MeV from QED correction.  $\rightarrow$  +3% above 4 MeV from using correct Z ("old way" use constant Z to fit all virtual branches) Activities was simulated by MCNP(Monte-Carlo N-Particle transport code) for Reactor Evolution.



after 12h, after 36h, accumulate. Time variation affect  $\pm$ 1%. Total error for  $^{235}$ U,  $^{239}$ Pu,  $^{241}$ Pu < 4% at 2-5 meV.



Using ENSDF, Pandemonium-corrected, JENDL and model. Compare with [Vogel, 1981](different nuclear database)  $\pm 10\%$ .  $^{238}$ U+n  $\rightarrow ^{239}$ U  $\xrightarrow{\rightarrow} ^{239}$ Np  $\xrightarrow{\rightarrow} ^{239}$ Pu  $^{239}$ Pu

# **Time evolution**



# n capture on fission product

 $\begin{array}{l} \mbox{Mainly on } ^{135}\mbox{Xe, very strong n-absorber.} \\ ^{135}\mbox{Xe} \rightarrow {}^{135}\mbox{Cs.} \\ \mbox{or } {}^{135}\mbox{Xe} + n \rightarrow {}^{136}\mbox{Xe(stable)} \end{array}$ 



The effect is minor. [Kipeikin, 2004]



Average  $N_{obs}/N_{pred}$  = 0.937±0.027(used to 0.979±0.029).  $\rightarrow$  a sterile neutrino?

- $\pm$  ~1% on  $e^-$  spectrum of  $^{235}$ U,  $^{239}$ Pu,  $^{241}$ Pu.
- +  $\pm$  ~10% on new/old calculation on  $^{238}\text{U's}~\bar{v_e}.$
- + 3% shift above 2 MeV.
- < 2 MeV spectrum uncheck.